## Antigravitation in higher dimensional General Relativity

Alexander K. Guts Omsk State University, Omsk, Russia E-mail: aguts@mail.ru

Let's consider 5-metrics

$$dS^{2} = \left[1 + \frac{1}{6}(\varkappa c^{2}\rho_{2}a - 2\Lambda_{1})ar^{2}\right]dx^{0^{2}} - \left[1 - \frac{(\varkappa c^{2}\rho_{2}a + \Lambda_{1})}{3} \cdot r^{2}a\right]^{-1}dr^{2} - r^{2}d\Omega^{2} - da^{2},$$
$$d\Omega^{2} = d\theta^{2} + \sin^{2}\theta d\varphi^{2},$$

where  $\varkappa, \rho_2, \Lambda_1 = const.$  Gravitational force, operating on a trial body, in 4-dimensional space-time  $\langle M_a^4, ds^2 \rangle = \langle (x^0, r, \theta, \varphi), dS^2 |_{a=const} \rangle$ , is possible to calculate on a formula from [Landau L., Lifshits E. Theory of field. Moscow, 1973. P.327]:

$$f_{\alpha} = \frac{mc^2}{\sqrt{1 - \frac{v^2}{c^2}}} \left\{ -\frac{\partial}{\partial x^{\alpha}} \ln \sqrt{g_{00}} + \sqrt{g_{00}} \left[ \frac{\partial}{\partial x^{\beta}} \left( \frac{g_{0\alpha}}{g_{00}} \right) - \frac{\partial}{\partial x^{\alpha}} \left( \frac{g_{0\beta}}{g_{00}} \right) \right] \frac{v^{\beta}}{c} \right\}.$$

We have

$$f_r = -\frac{mc^2}{6\sqrt{1 - v^2/c^2}} \frac{\left[\varkappa c^2 \rho_2 a - 2\Lambda_1\right] ar}{\left[1 + \frac{1}{6}(\varkappa c^2 \rho_2 a - 2\Lambda_1) ar^2\right]}, \ f_{\varphi} = f_{\theta} = 0.$$

In this case, it is obvious that it is possible to find the functions  $\Lambda = \Lambda_1 a$ ,  $\rho = \rho_2 a^2$  so that  $\rho_2 > 0$ , and  $f_r$  changes a sign in a = 0 and  $a = 2\Lambda_1/(\varkappa c^2 \rho_2)$  in extensive spatial area with radius  $r < c\sqrt{6\varkappa\rho_2}/|\Lambda_1|$  (Inequality is received as a condition of positivity of a denominator in a formula for  $f_r$  for every a.), i.e. the attraction to the center of r = 0 is replaced by the repulsion from center r = 0.

Transition through a = 0 changes a sign of "the cosmological constant"  $\Lambda$ , and observable change of gravitation on antigravitation can be regarded as manifestation of the cosmological repulsion. But upon transition through  $a = 2\Lambda_1/(\varkappa c^2 \rho_2)$  "the cosmological constant" keeps a sign, and it means that we have other type of antigravitation.

If  $r > c\sqrt{6\varkappa\rho_2}/|\Lambda_1|$ , then denominator of  $f_r$  is remained positive under

$$a > a_{+}(r) = \frac{1}{\varkappa c^{2} \rho_{2}} [\Lambda_{1} + \sqrt{\Lambda_{1}^{2} - (6\varkappa c^{2} \rho_{2}/r^{2})}] \text{ or } a < a_{-}(r) = \frac{1}{\varkappa c^{2} \rho_{2}} [\Lambda_{1} - \sqrt{\Lambda_{1}^{2} - (6\varkappa c^{2} \rho_{2}/r^{2})}].$$

If  $\Lambda_1 > 0$ , then we have  $a_+(r) < 2\Lambda_1/(\varkappa c^2 \rho_2)$ . Hence, for every  $r > c\sqrt{6\varkappa \rho_2}/\Lambda_1$  when the parameter a is changed in some small interval  $(2\Lambda_1/(\varkappa c^2 \rho_2) - \varepsilon(r), 2\Lambda_1/(\varkappa c^2 \rho_2) + \varepsilon(r))$ , the function  $f_r$  changes a sign, i.e. the attraction is replaced by the repulsion. Under  $\Lambda_1 < 0$  we have  $2\Lambda_1/(\varkappa c^2 \rho_2) < a_-(r)$ . Hence, for every  $r > c\sqrt{6\varkappa \rho_2}/\Lambda_1$  gravitational force  $f_r$  changes sign, when the parameter a is changed in the same interval.

Thus, when we are moving in 5-dimensional bulk, i.e. when a is changed, the geometry of 4-brane  $M_a^4$  is changed so that gravitation (attraction) is replaced with antigravitation (repulsion).